

国家天文台青年学术交流

云核速度弥散探测分子云

钱磊

国家天文台

FAST运行和发展中心

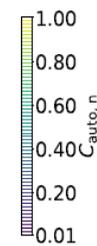
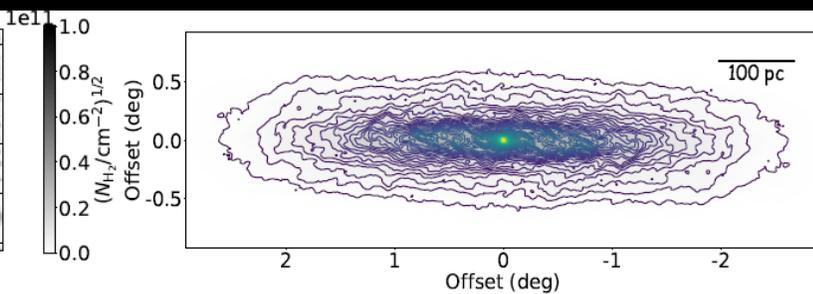
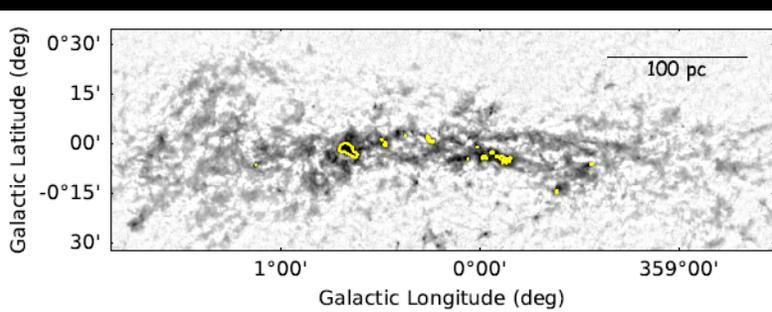
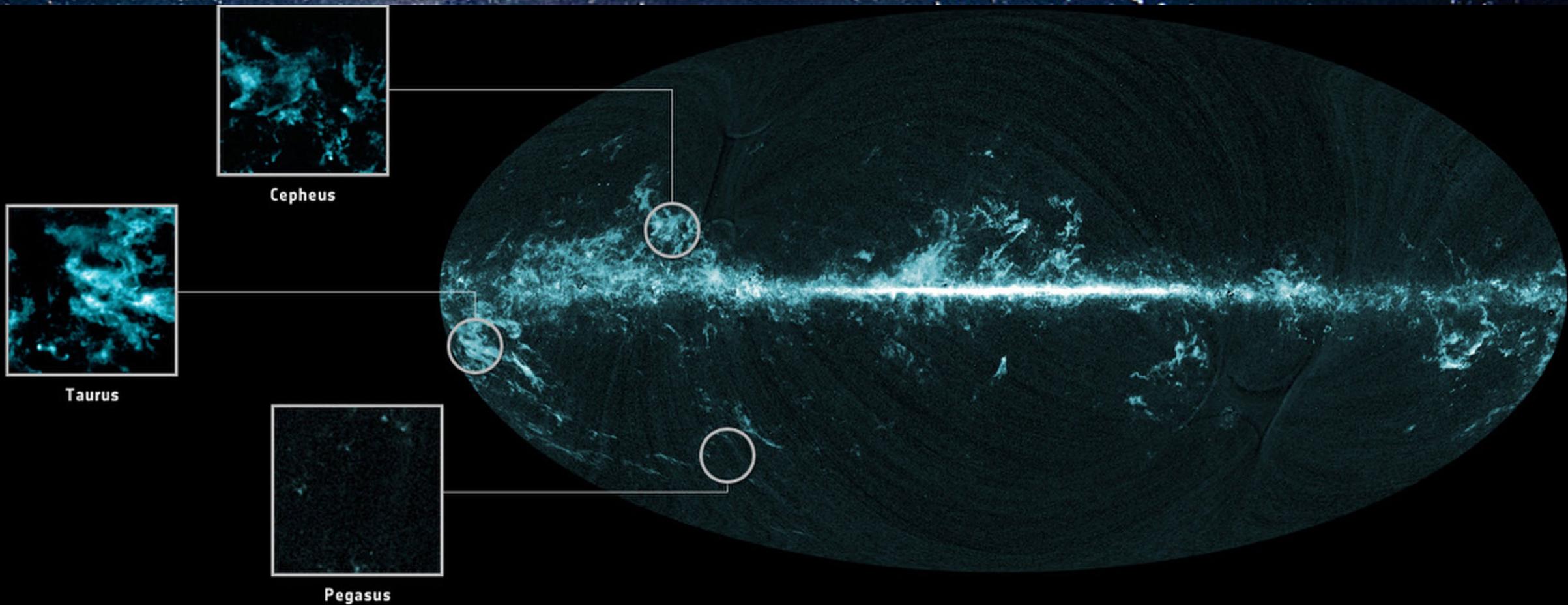
2022年11月 · 北京

- 01 分子云结构和湍流运动
- 02 云核速度弥散方法
- 03 分子云厚度和湍流耗散率
- 04 总结和展望

01

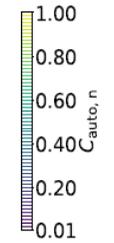
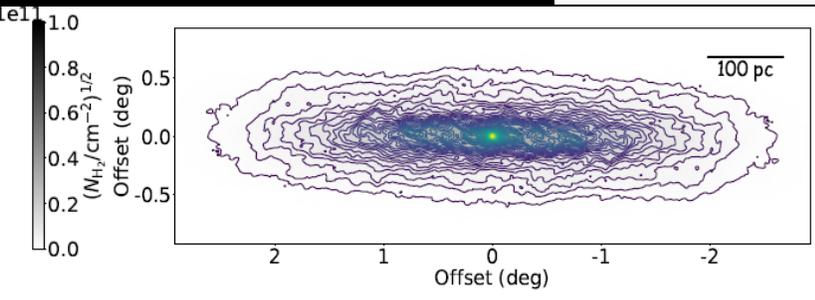
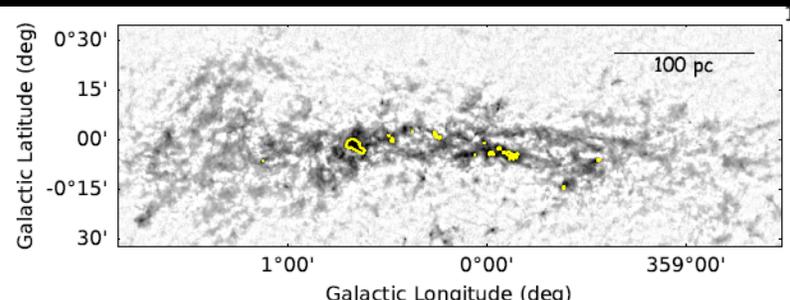
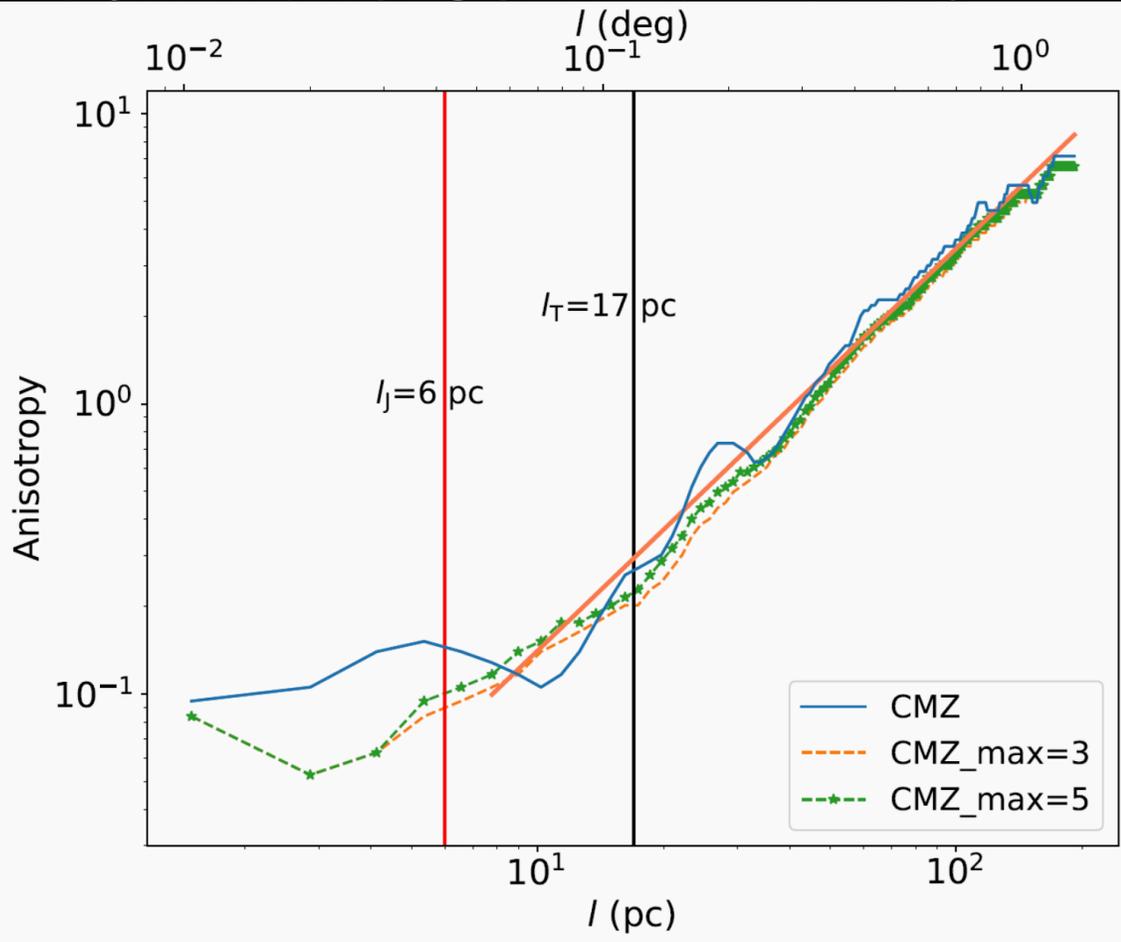
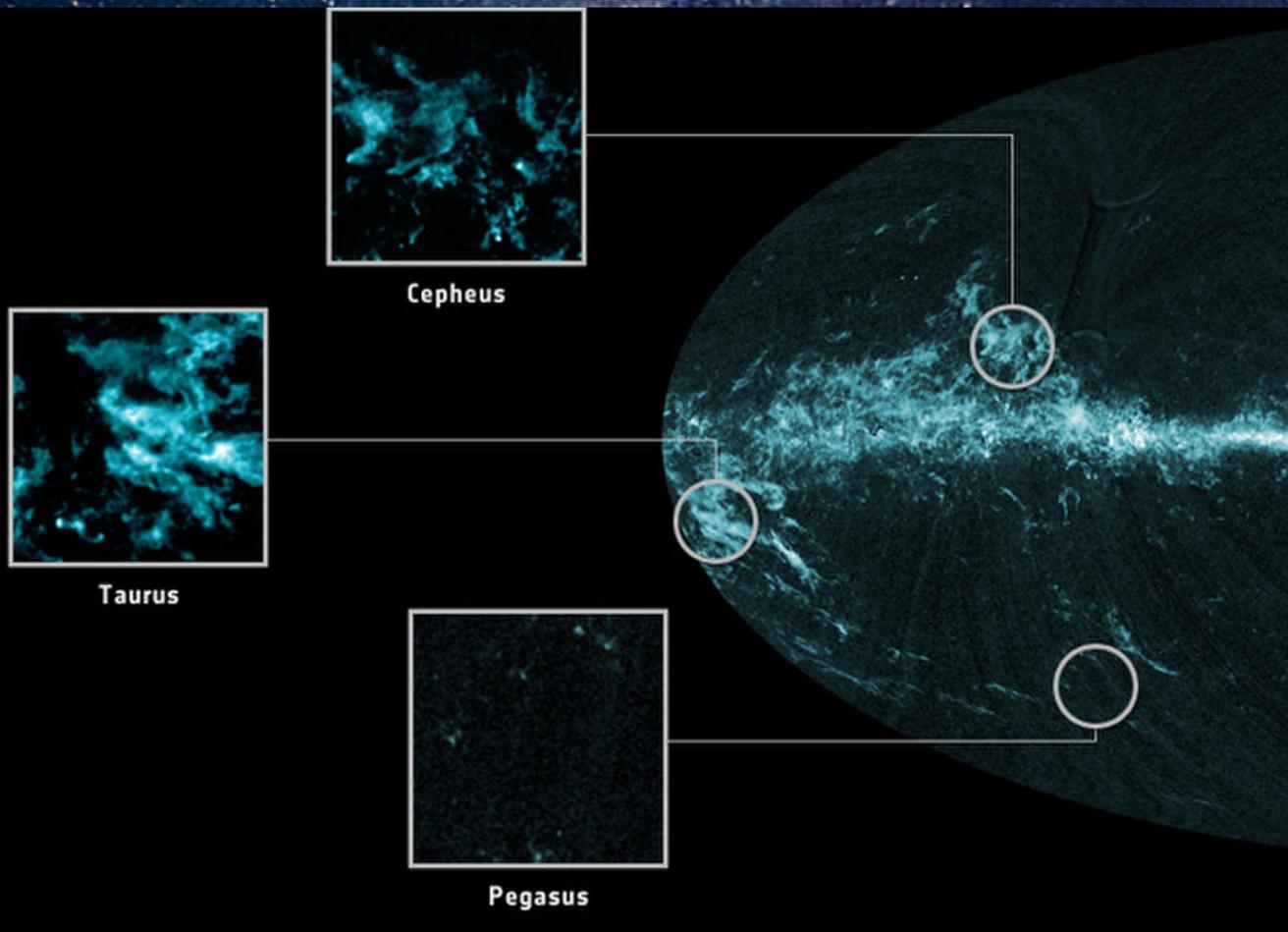
分子云结构和湍流运动

分子云结构和湍流运动



Cai, Li, Qian 2021, RAA, 21, 194

分子云结构和湍流运动



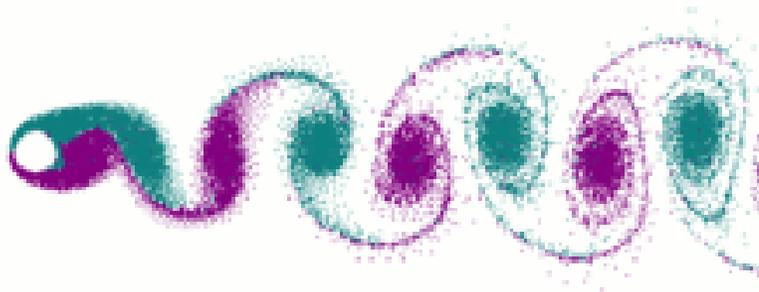
Cai, Li, Qian 2021, RAA, 21, 194

分子云结构和湍流运动

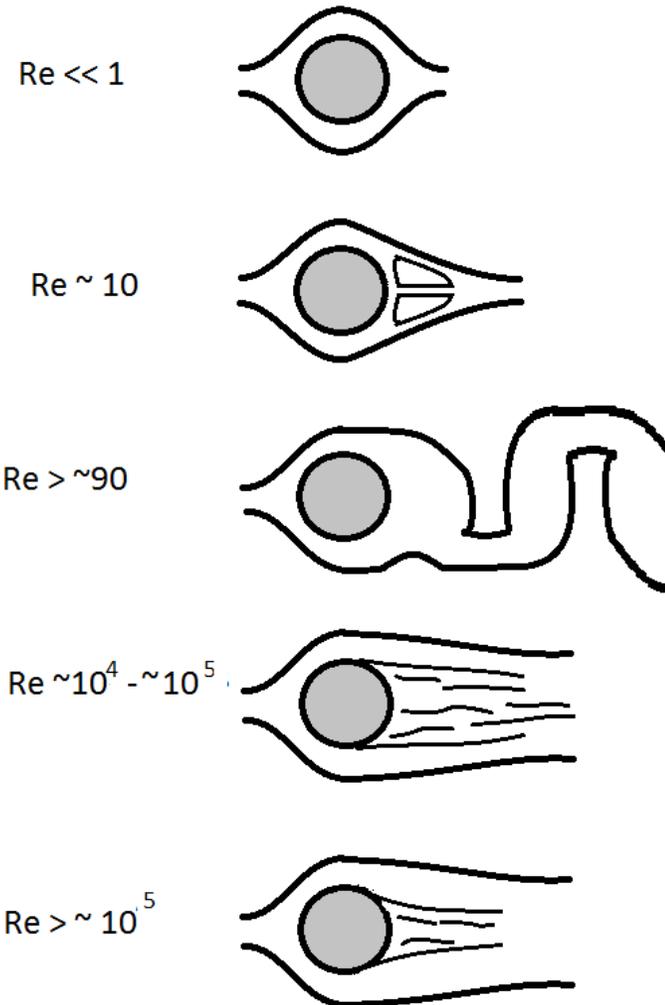
雷诺数

Reynolds Number

$$Re = \frac{\text{inertial forces}}{\text{viscous forces}} = \frac{\rho v L}{\mu} = \frac{v L}{\nu}$$



——Wikipedia

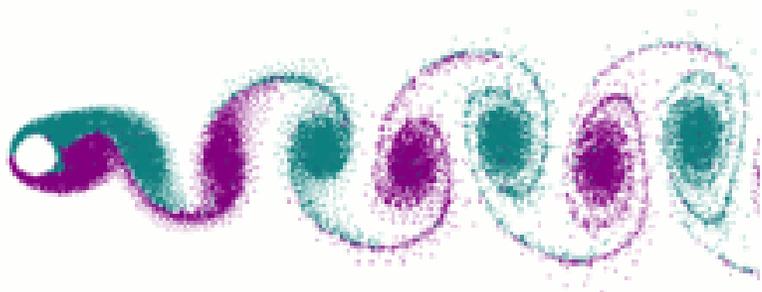


分子云结构和湍流运动

雷诺数

Reynolds Number

$$Re = \frac{\text{inertial forces}}{\text{viscous forces}} = \frac{\rho v L}{\mu} = \frac{v L}{\nu}$$

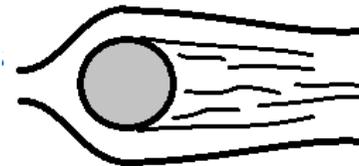


——Wikipedia

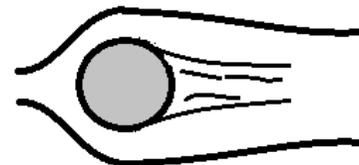
分子云

Molecular Clouds

$Re \sim 10^4 - \sim 10^5$



$Re > \sim 10^5$



$$\rho = 1.67 \times 10^{-22} \text{ g/cm}^3$$

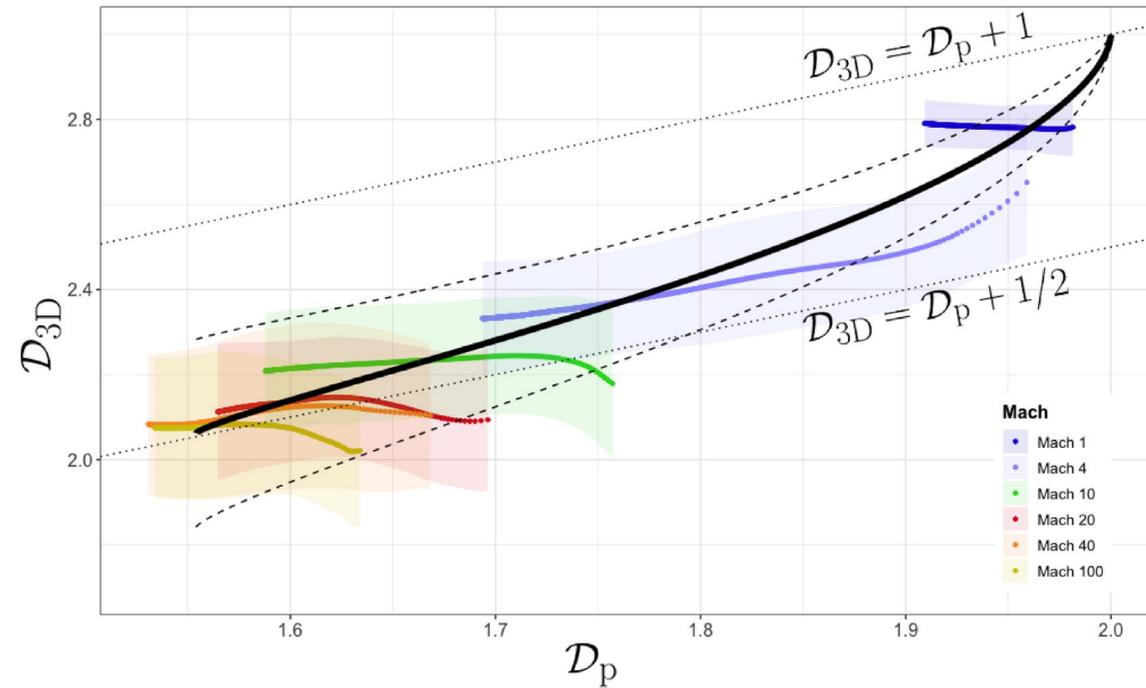
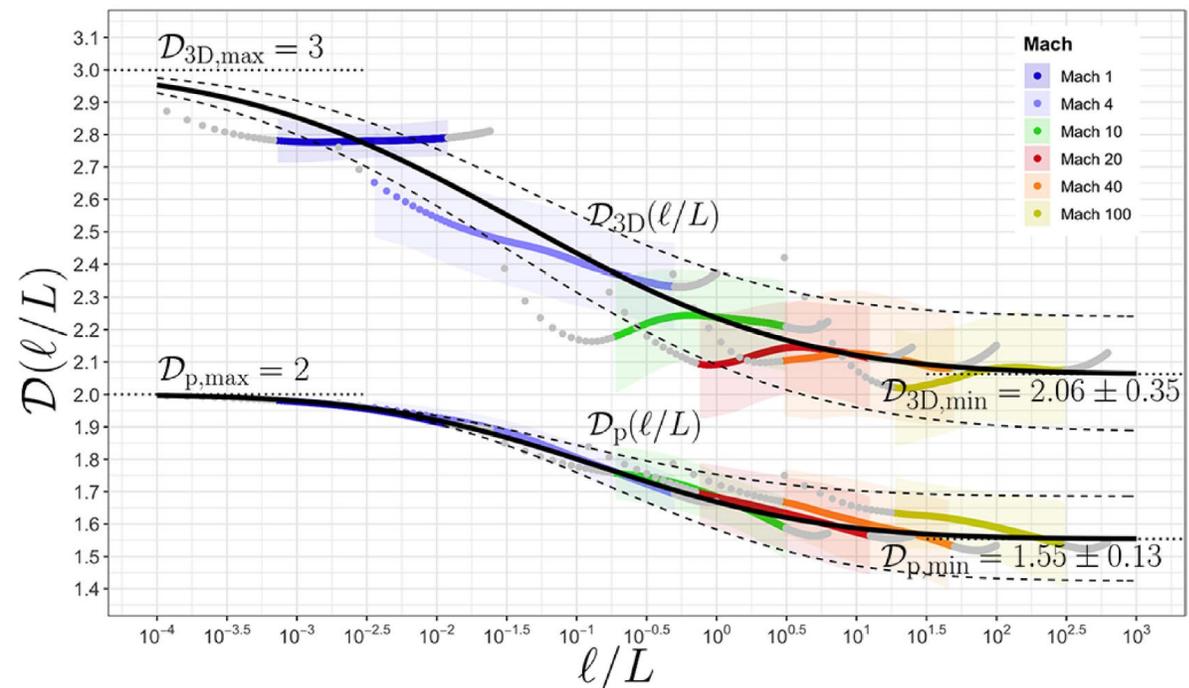
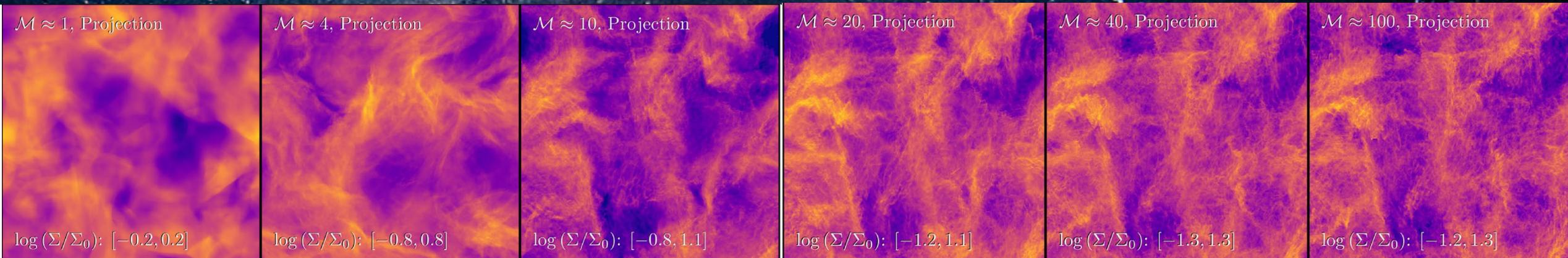
$$v = 10 \text{ km/s}$$

$$L = 10 \text{ pc}$$

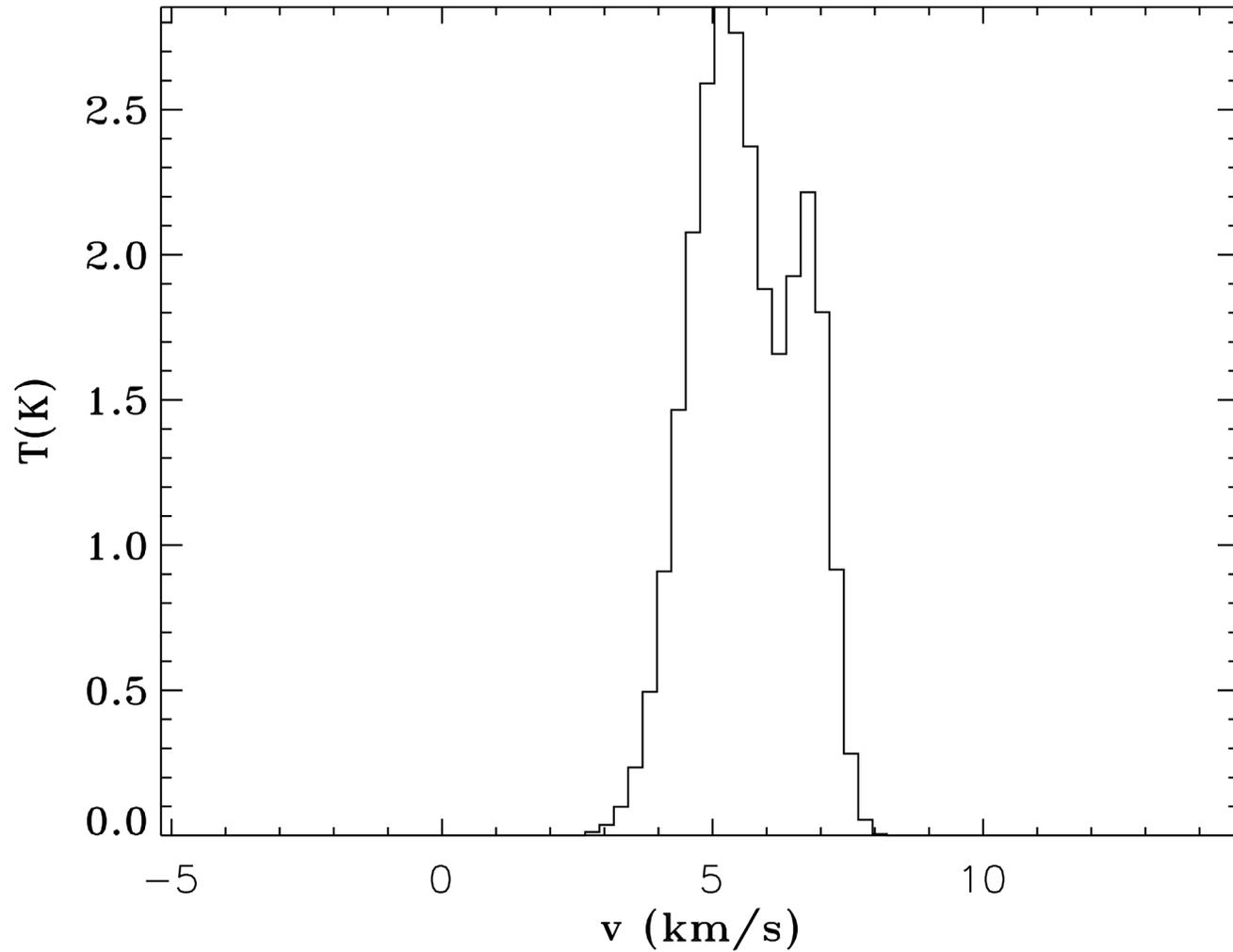
$$\mu = 2.5 \times 10^{-7} \text{ kg / (m s)}$$

$$Re \sim 10^9$$

分子云结构和湍流运动



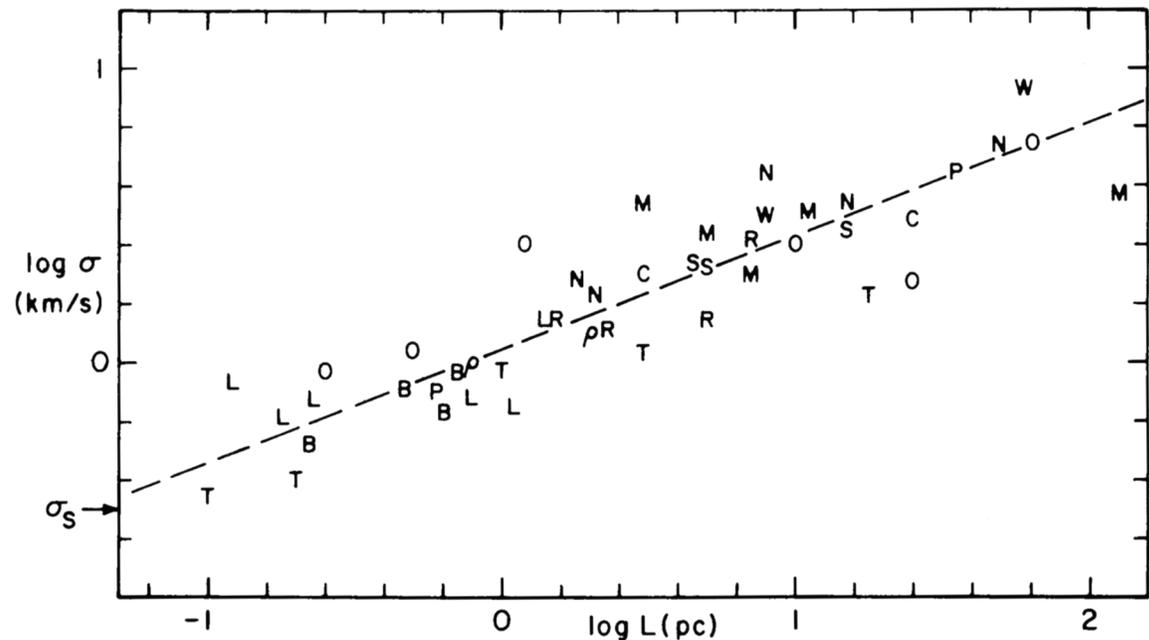
分子云结构和湍流运动



$$(kT/m)^{1/2} \\ \sim 0.05 \left(\frac{T}{10\text{K}} \right)^{0.5} \left(\frac{m}{m_{\text{CO}}} \right)^{-0.5} \text{ km s}^{-1}$$

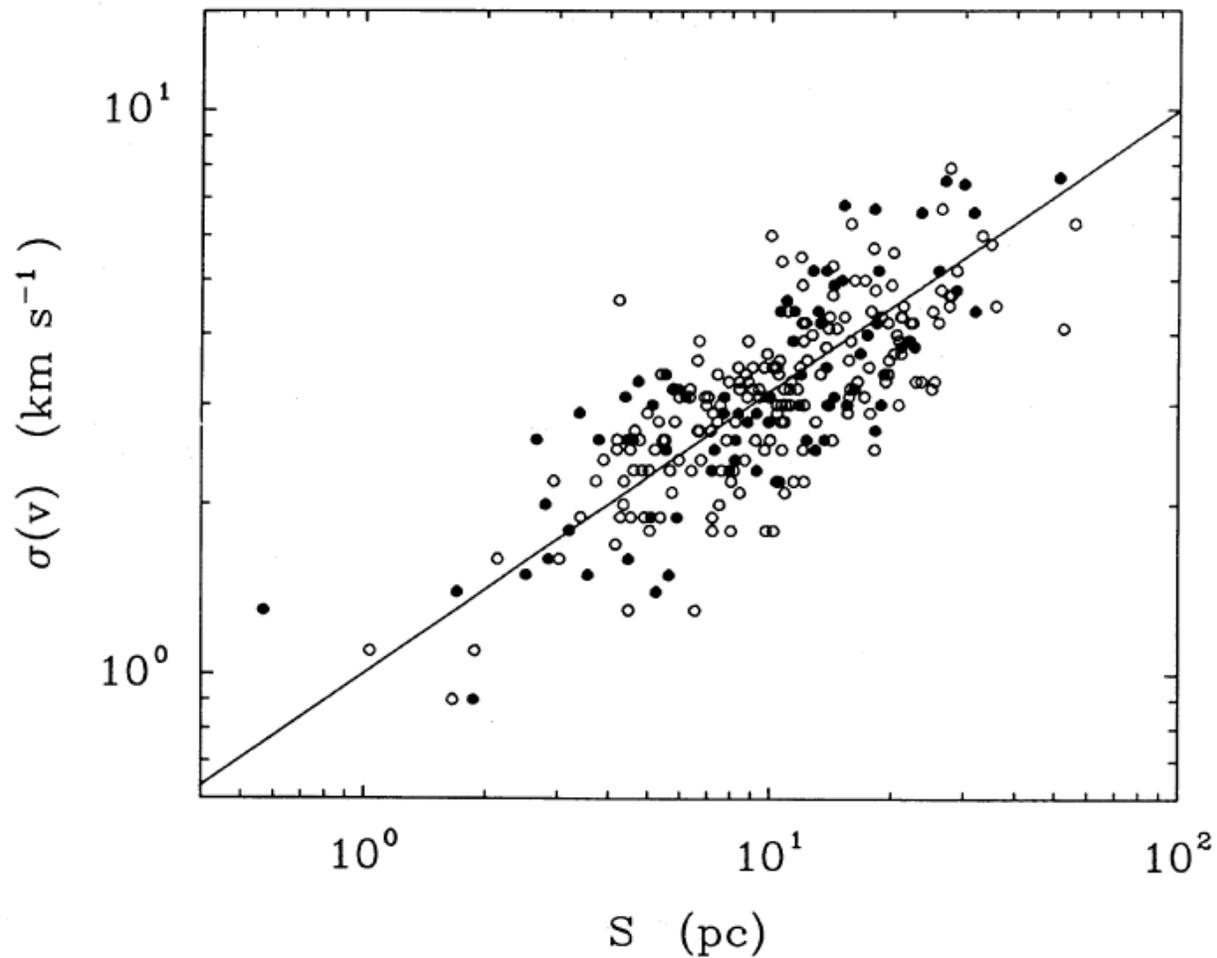
Qian, Li, Goldsmith, 2012, ApJ, 760, 145

分子云结构和湍流运动



$\sigma \propto S^{0.38}$ (Larson 关系) \rightarrow 不可压缩湍流

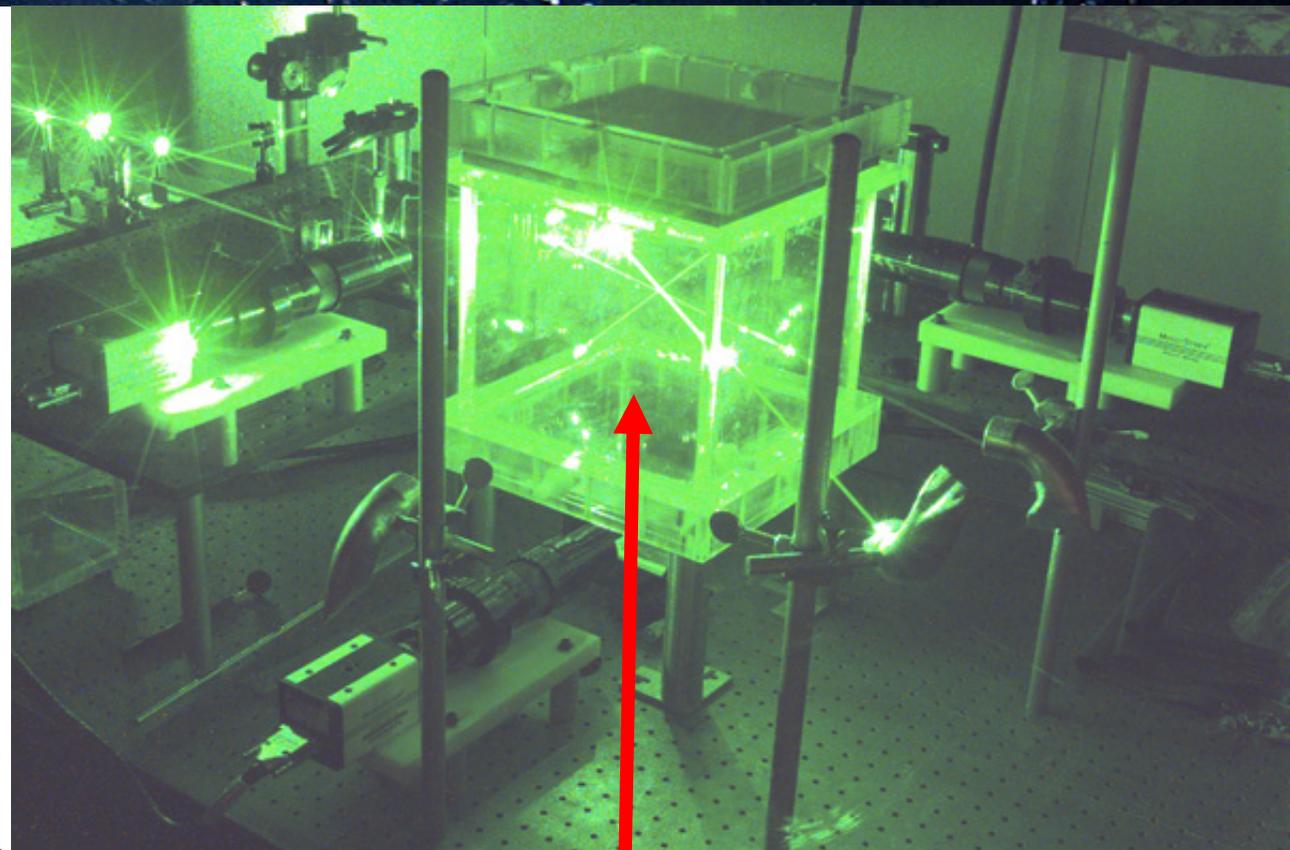
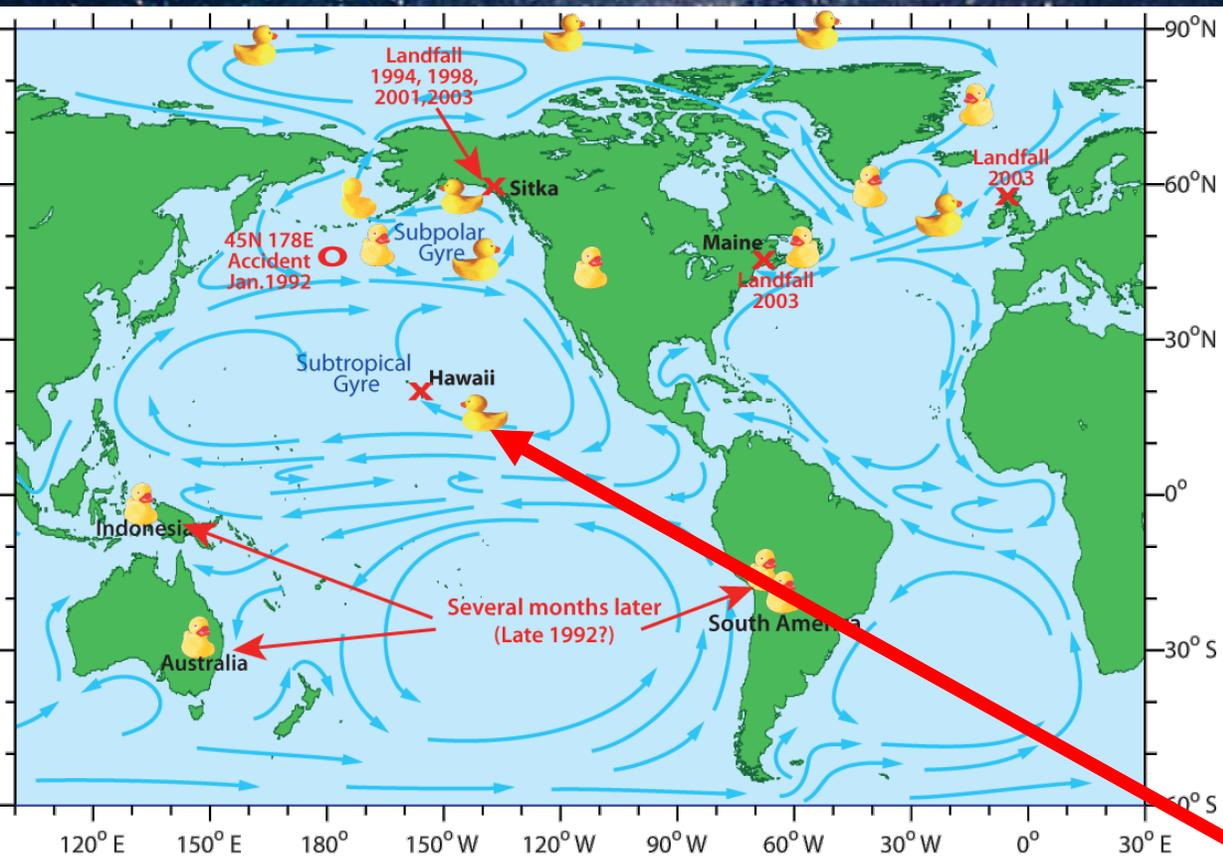
Larson, 1981, MNRAS, 194, 809



$\sigma \propto S^{0.5}$ (修正的Larson 关系) \rightarrow 可压缩湍流

Solomon et al., 1987, ApJ, 319, 730

分子云结构和湍流运动



流动的实验测量

示踪粒子



分子云结构和湍流运动

Genus statistics

Dendrograms

Modified velocity centroids

Spatial power spectrum

Bispectrum and bicoherence

Velocity coordinate spectrum (VCS)

Velocity channel analysis (VCA)

Delta variance

Wavelet transform

Principal component analysis eigenvalues

Spectral correlation function

Intensity probability density function

Higher order statistical moments

Cramer statistic

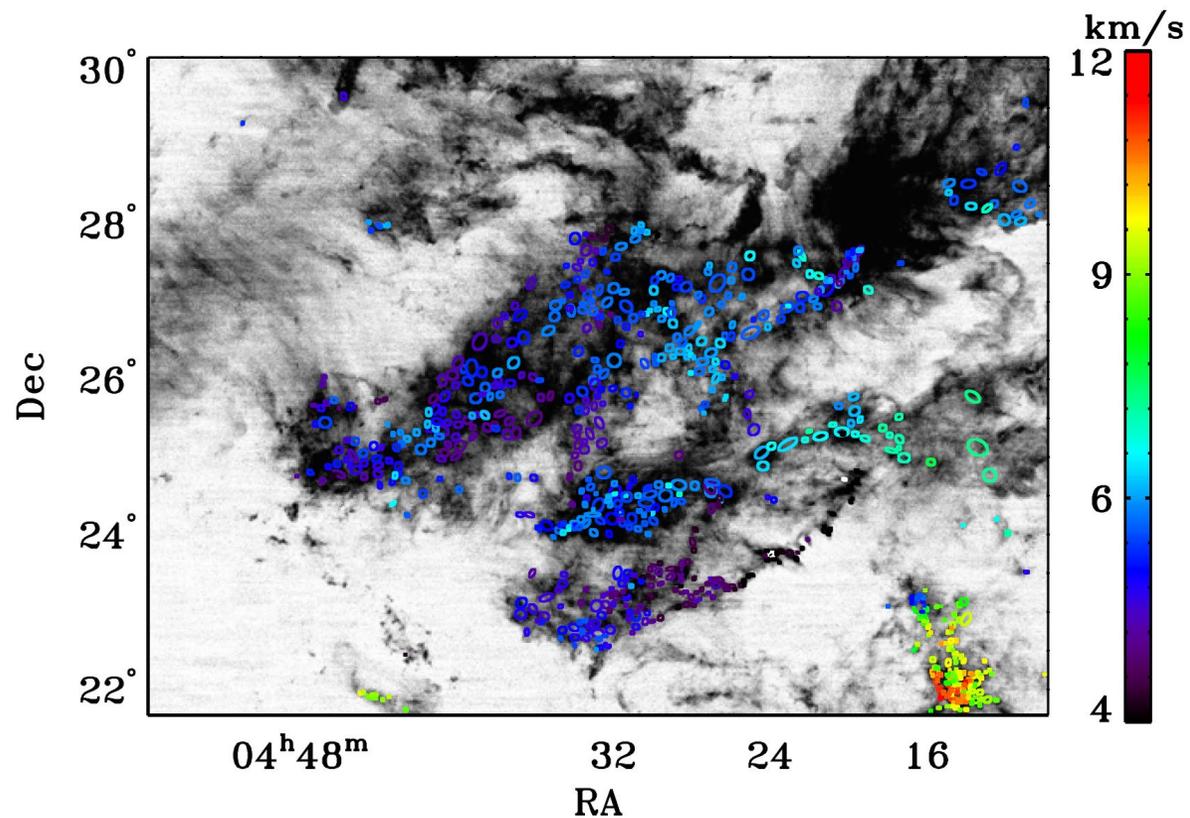
Statistic	Noiseless		Free-fall		Added noise	
	Noise <i>p</i> -value	Signal <i>p</i> -value	Noise <i>p</i> -value	Signal <i>p</i> -value	Noise <i>p</i> -value	Signal <i>p</i> -value
VCS	0	1	0	1	0.0751	0
VCS Large Scale	0	1	0	1	0.0689	0
SCF	0	1	0	1	0	1
Skewness	0	1	0	1	0	1
VCS Small Scale	0.0153	1	0.1534	0.9968	–	–
Del. Var. Centroid Curve	0.0001	1	0.0002	0.0147	0	0.9997
Kurtosis	0	1	0.0002	0.0001	0	0.2387
VCA	0	0.5308	0	0.2985	0	0.4956
Bispectrum	0.2442	0.3581	0	1	0.0993	0.3777
PDF Lognormal	0.0229	0.1963	0.0224	0.1306	0.0047	0.0012
Cramer	0	0.0299	0	0.0186	0	0.0272
Del. Var. Centroid Slope	0.0009	0.0412	0.0002	0.0056	0.0002	0.0961
MVC	0.0001	0.0008	0.0382	0	0.0003	0.0074
Spatial Power Spec.	0.0001	0.0008	0.0384	0	0.0003	0.0090
Genus	0.0051	0.0001	0.0109	0	0.0183	0.0004
Dendro. Num.	0.0992	0	0.0001	0.1830	0.0418	0.0007
PCA Eigenvalues	0.0060	0	0	0.6199	0.0404	0
Wavelet	0.0097	0	0.0870	0	0.0328	0
Del. Var. Curve	0.0315	0	0.0076	0	0.0848	0
Del. Var. Slope	0.0451	0	0.0044	0	0.0847	0
Dendro. Hist.	0.0122	0	0.0233	0	0.0083	0

[Koch et al., 2017, MNRAS, 471, 1506](#)

02

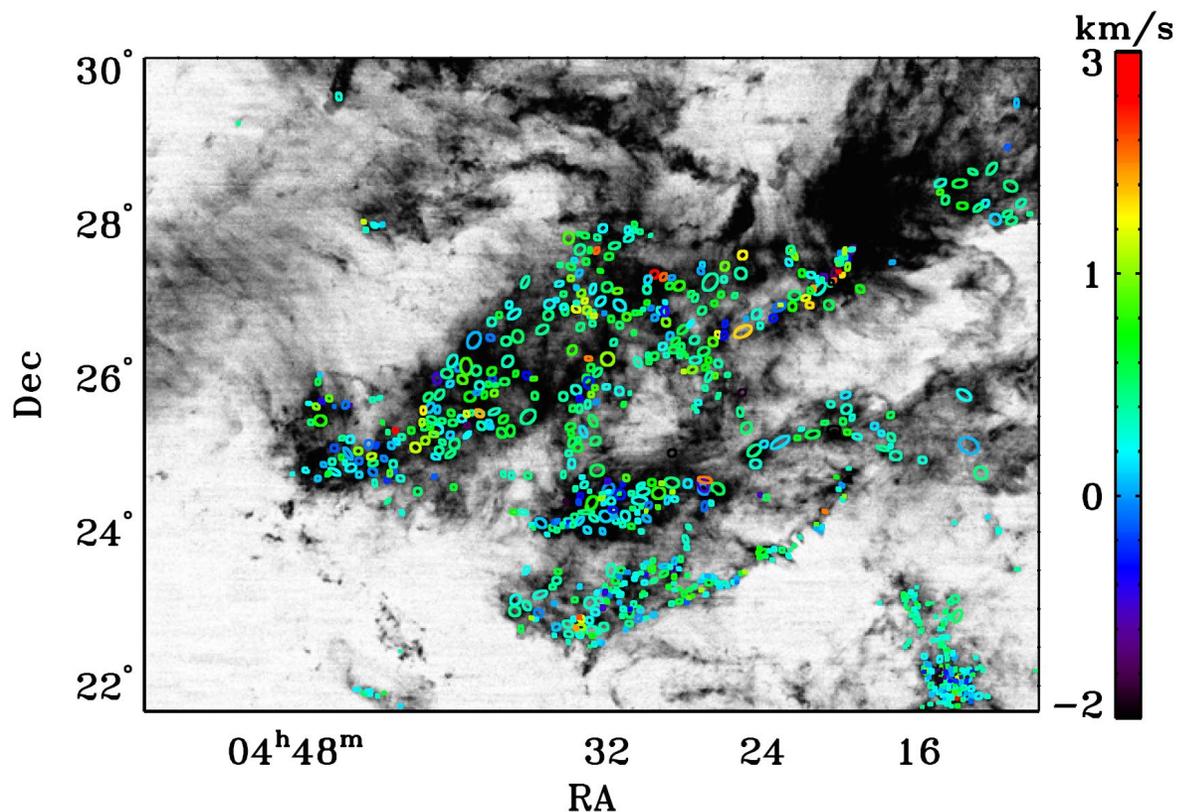
云核速度弥散方法

云核速度弥散方法



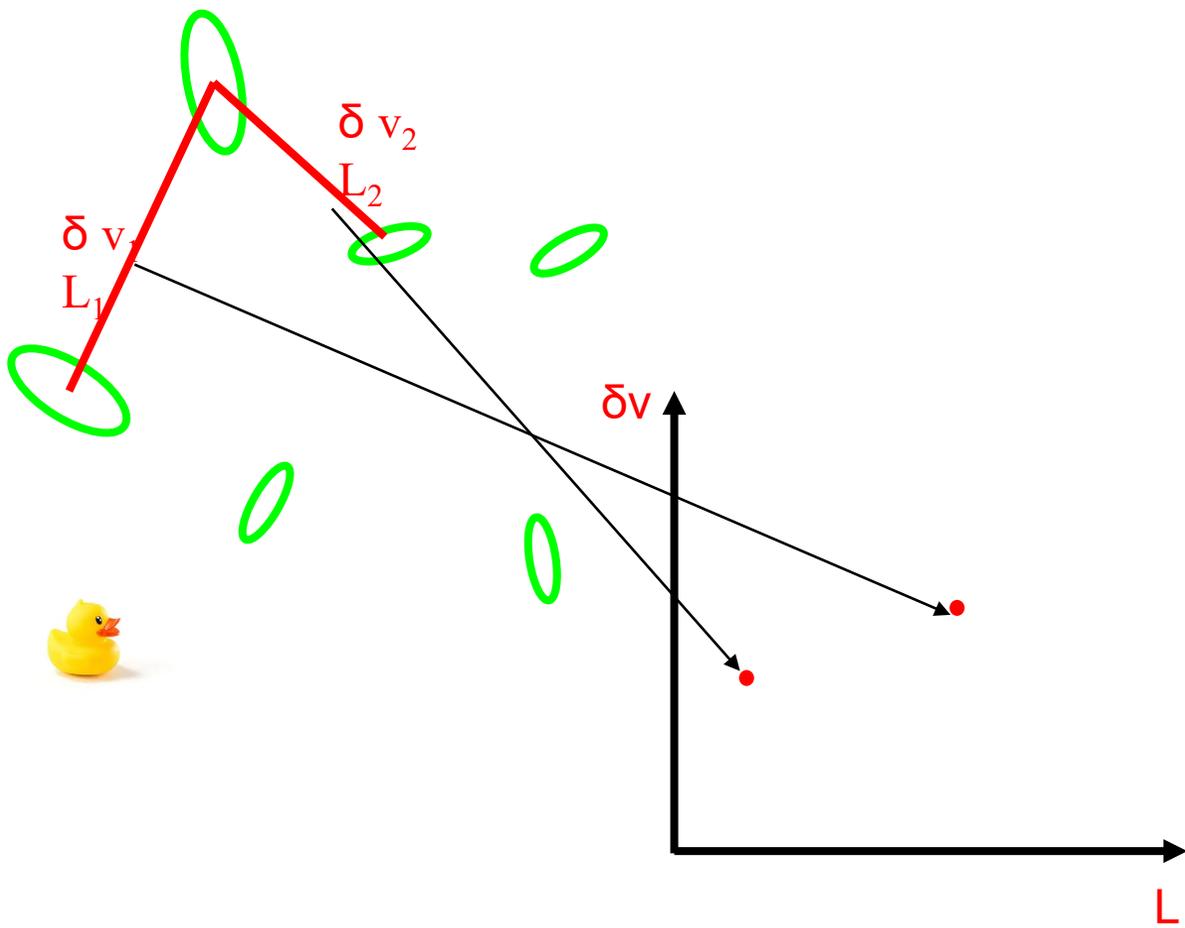
¹³CO云核中心速度

Qian, Li, Goldsmith, 2012, ApJ, 760, 145

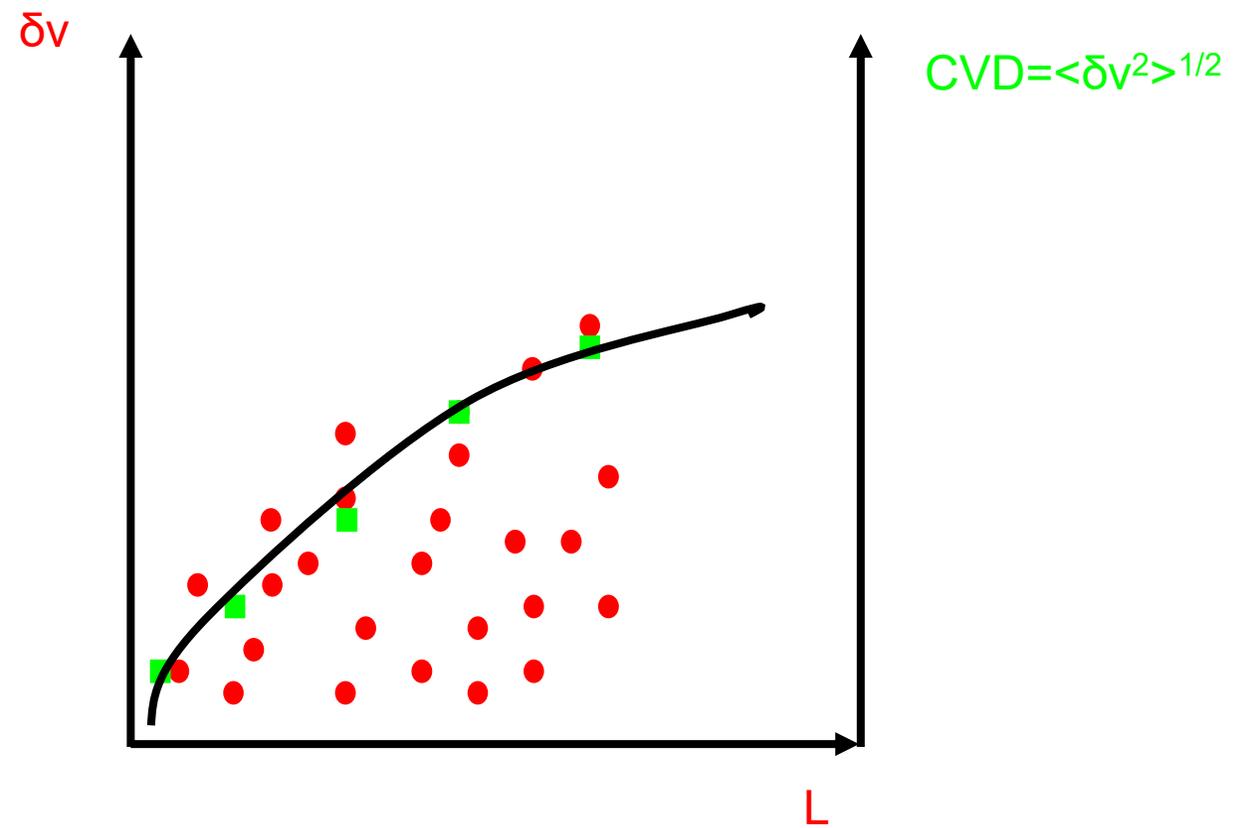


¹³CO云核中心速度与
¹²CO平均速度的差

云核速度弥散方法

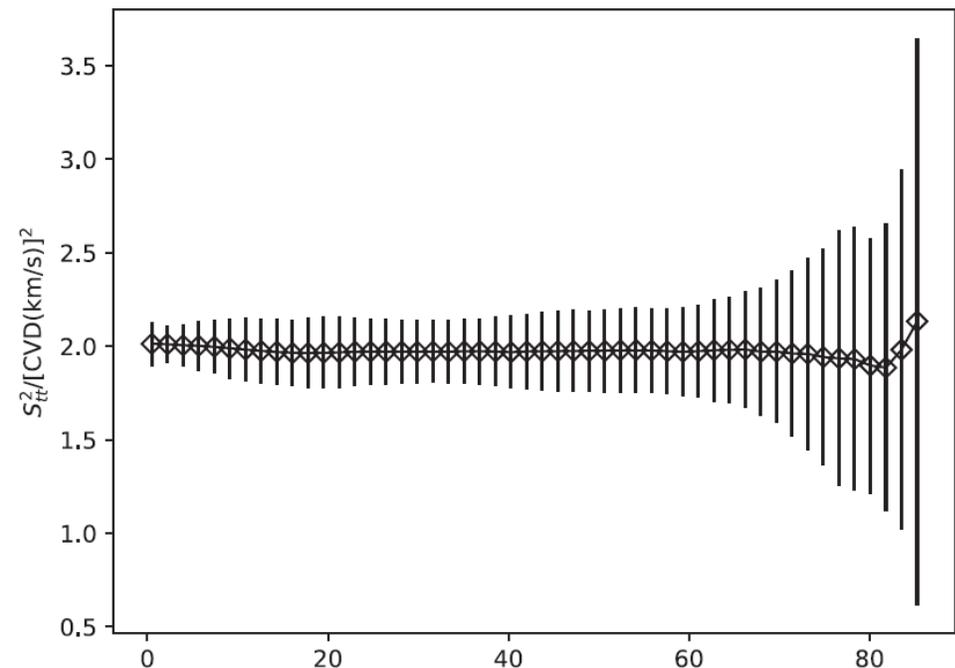
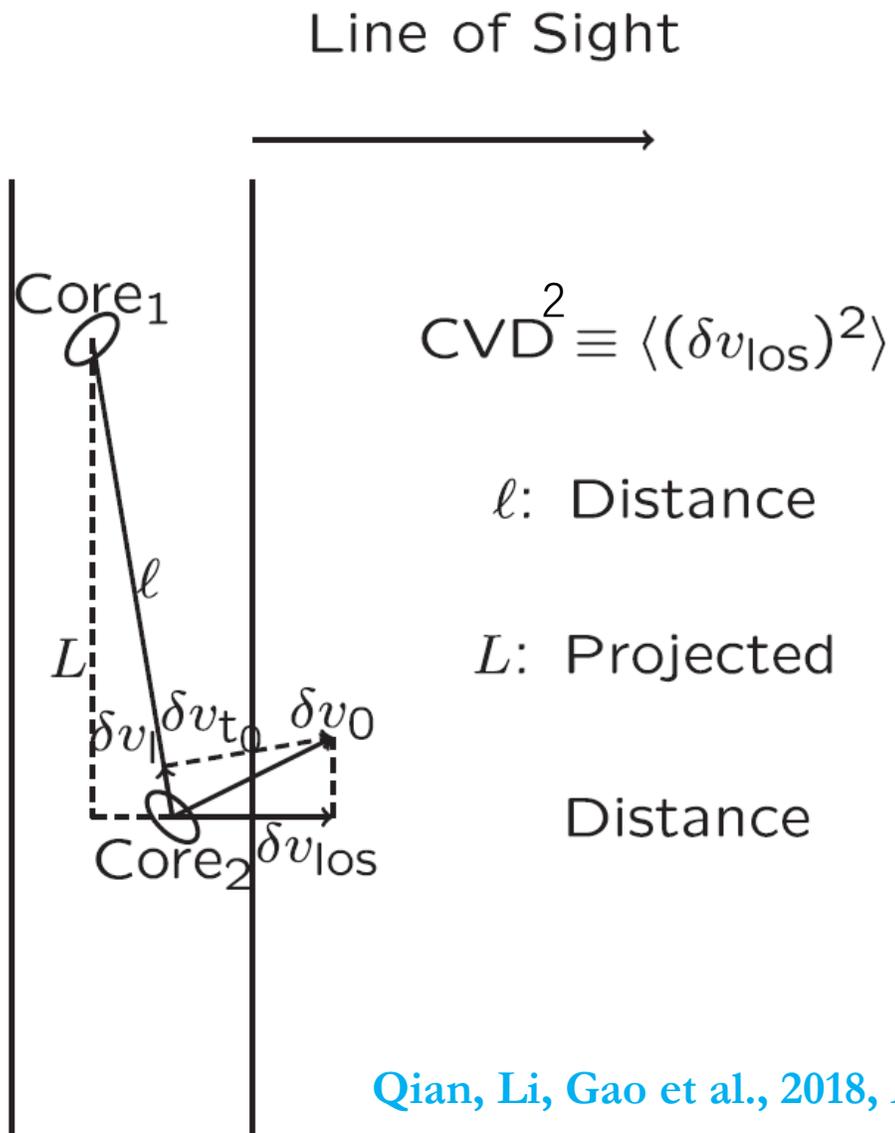


$$S_{tt}^2(l) = \langle [u_t(M') - u_t(M)]^2 \rangle$$



$$CVD^2 \equiv \langle \delta v^2 \rangle \approx \frac{1}{2} S_{tt}^2$$

云核速度弥散方法



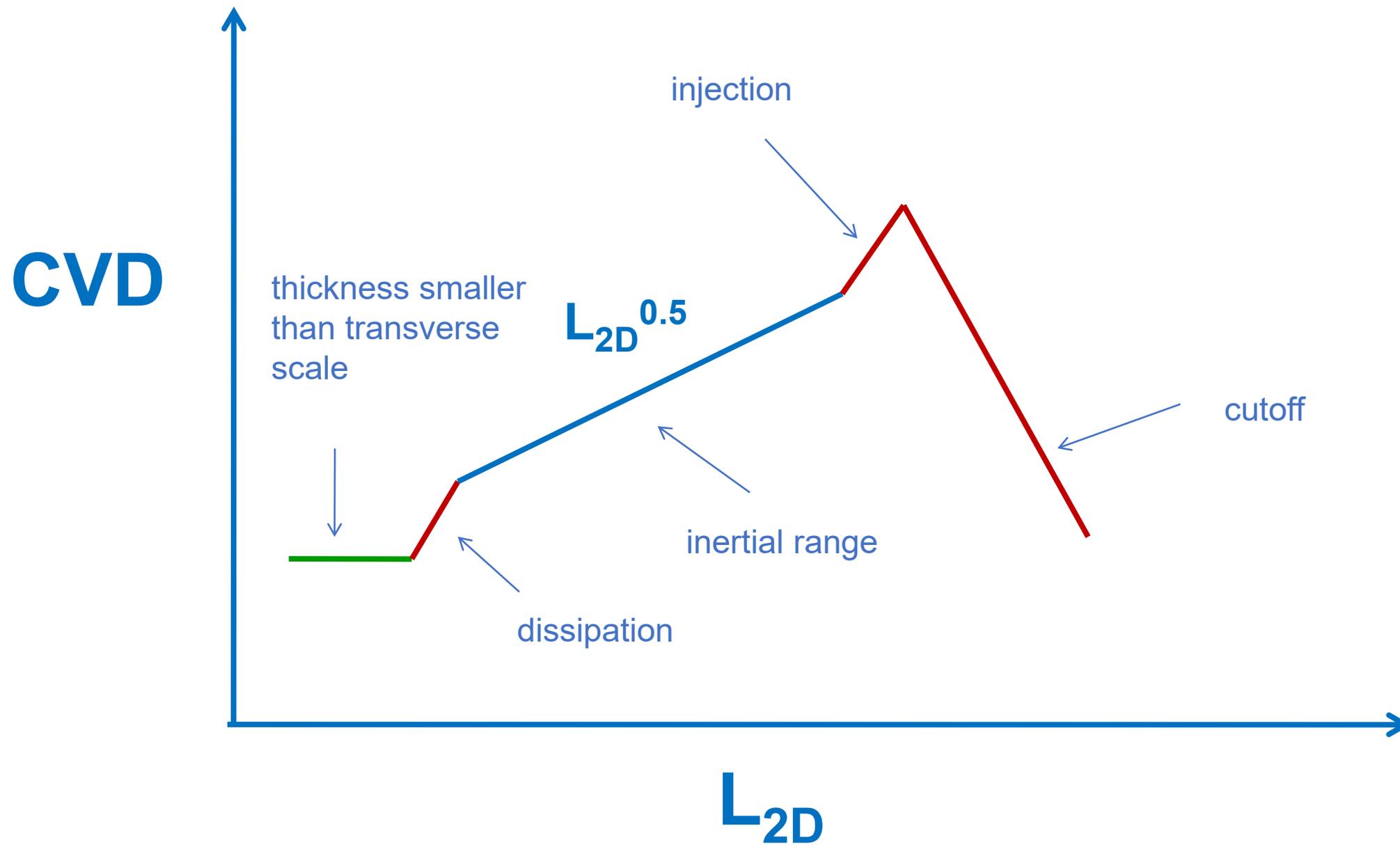
$$S_{tt}^2 = \frac{8}{3} C_{12} \epsilon^{2/3} l^{2/3}$$

$$CVD^2 \equiv \langle \delta v_{los}^2 \rangle = \frac{1}{2} f S_{tt}^2$$

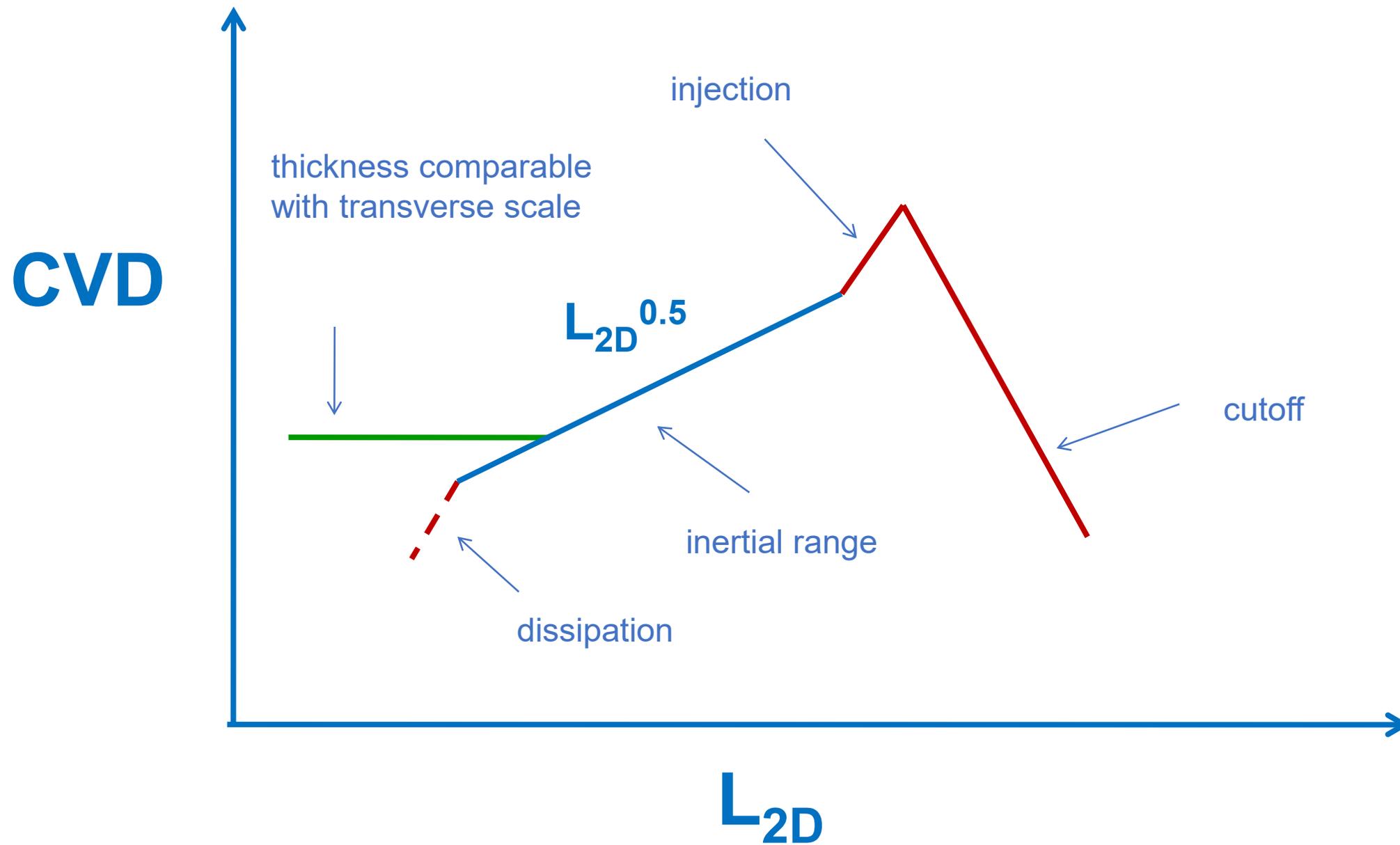
$$\epsilon = \frac{1}{L} \left(\frac{CVD^2}{\frac{4}{3} C_{12} f} \right)^{3/2}$$

Qian, Li, Gao et al., 2018, ApJ, 864, 116

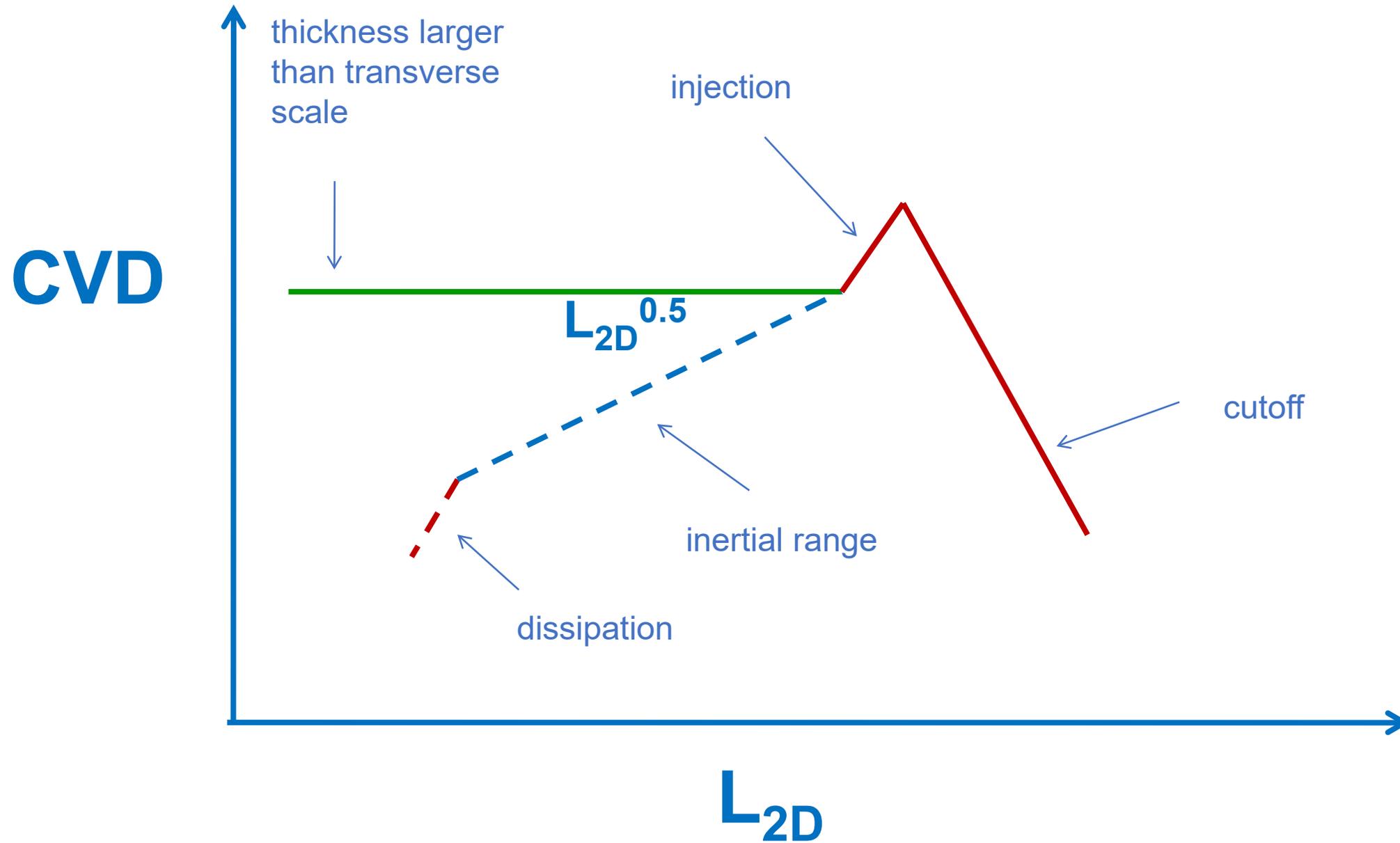
云核速度弥散方法



云核速度弥散方法



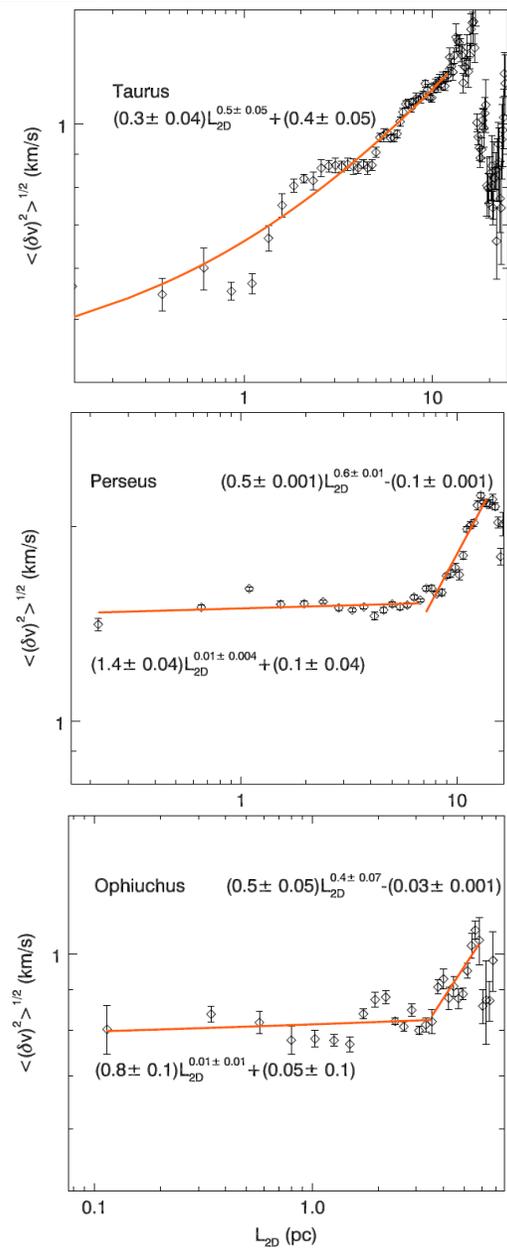
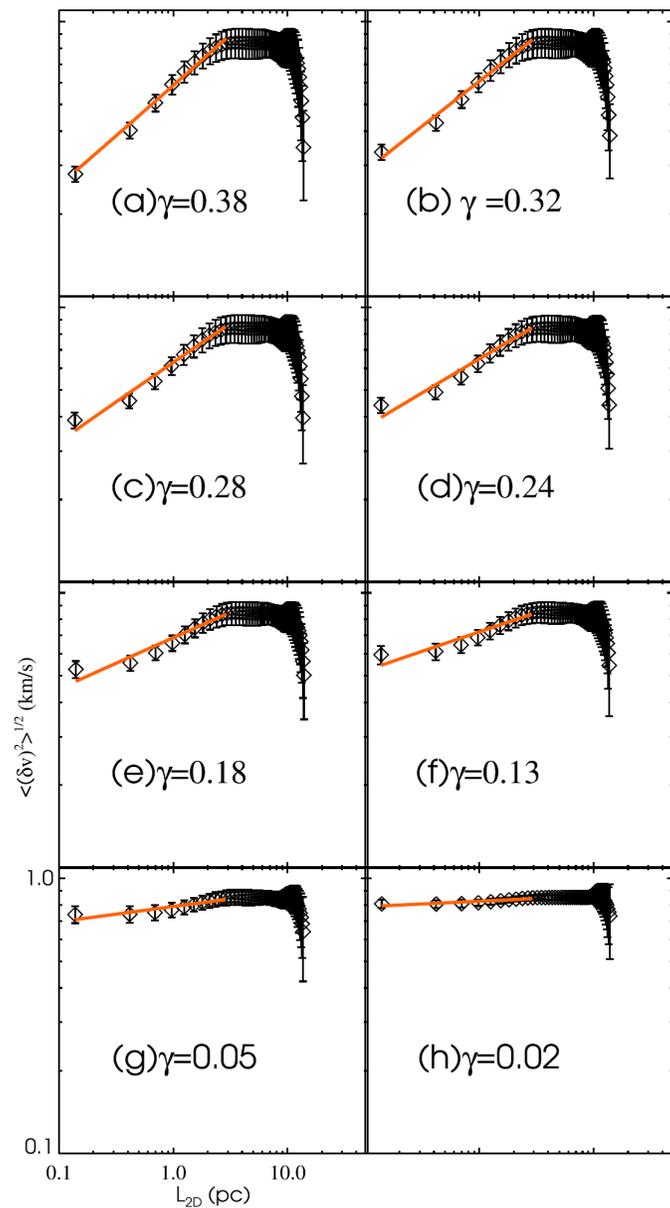
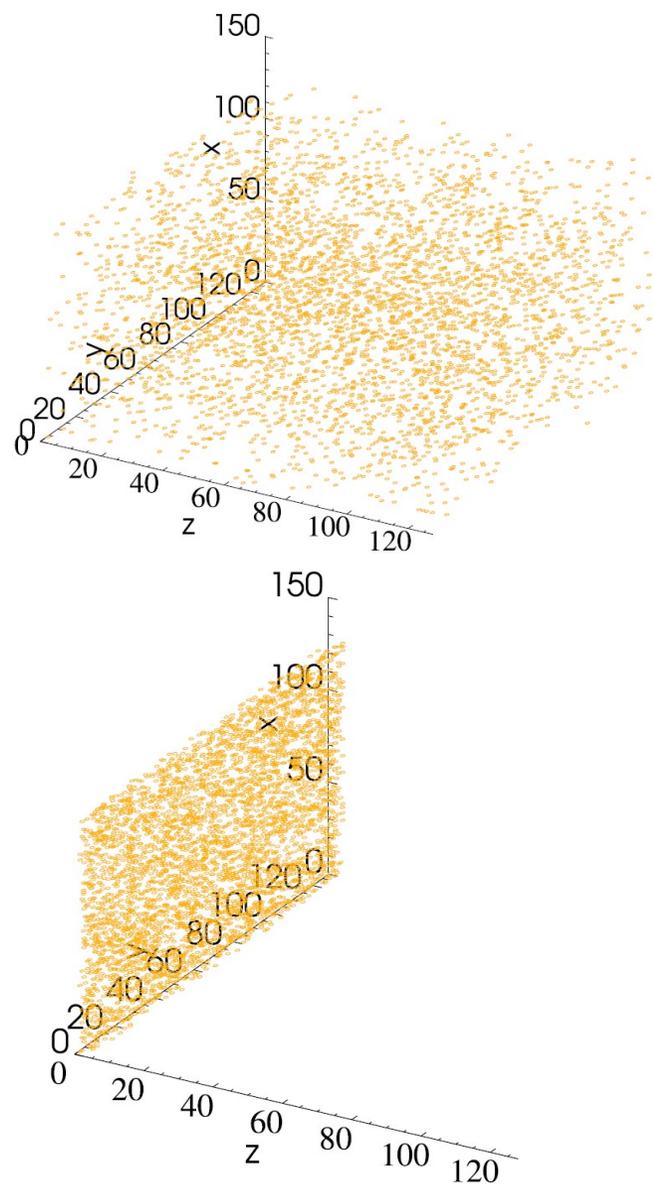
云核速度弥散方法



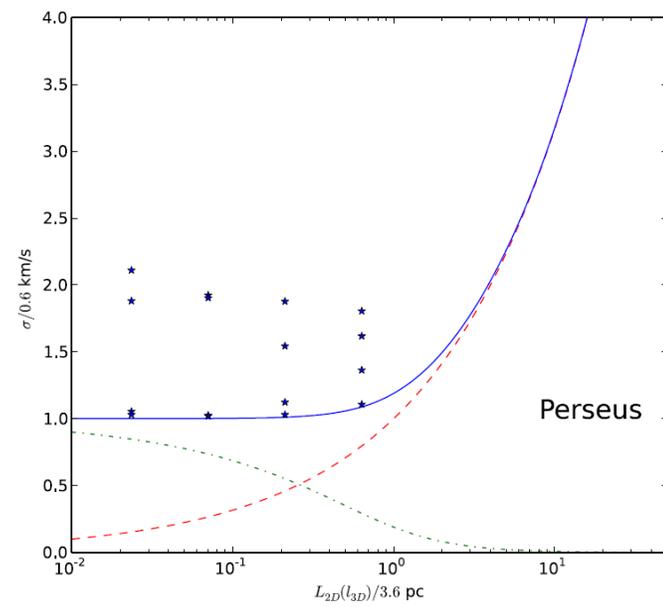
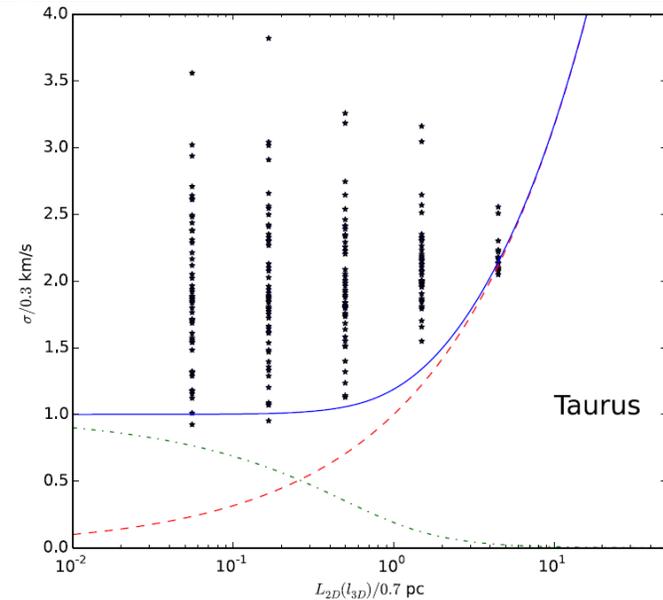
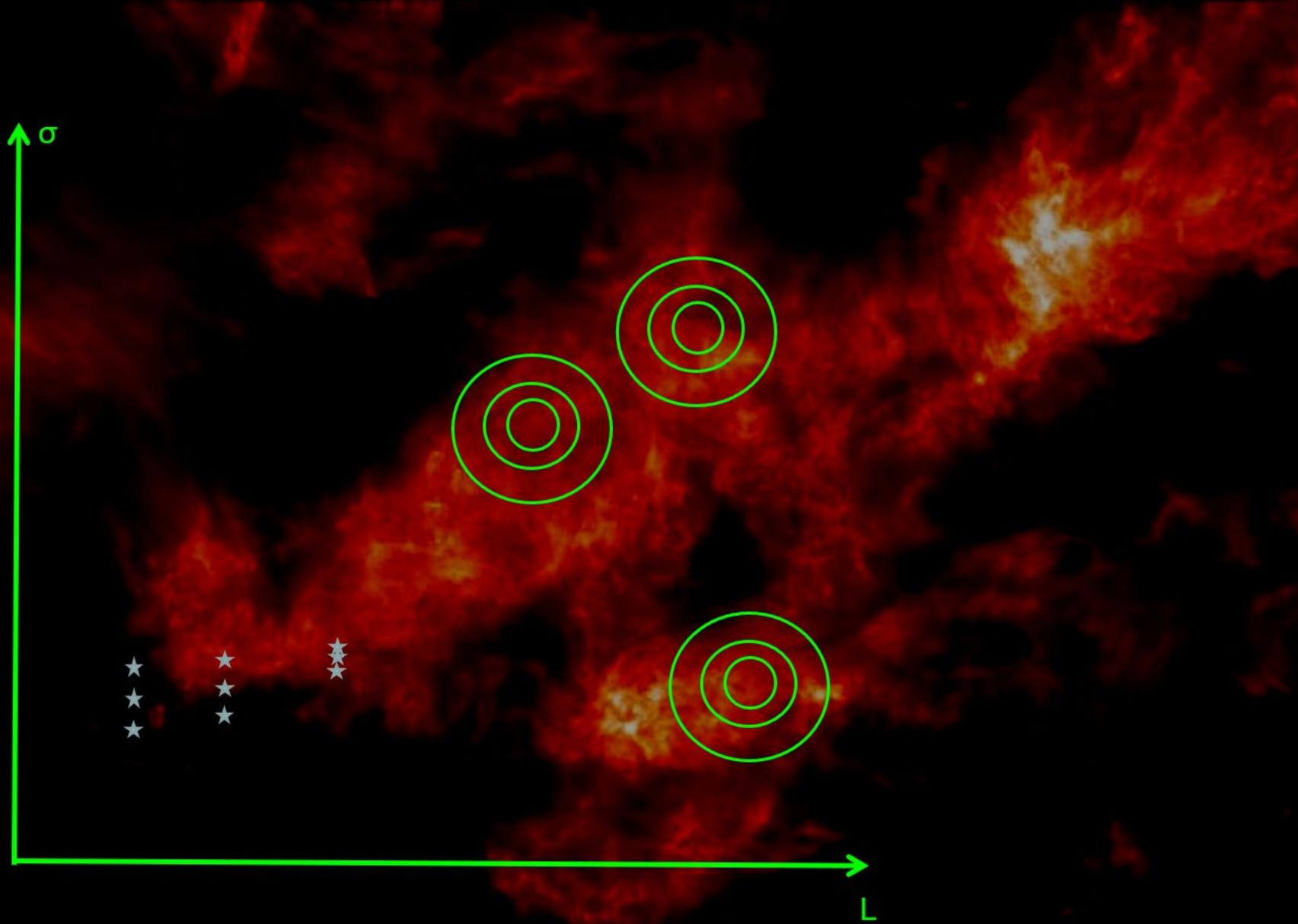
03

分子云厚度和湍流耗散率

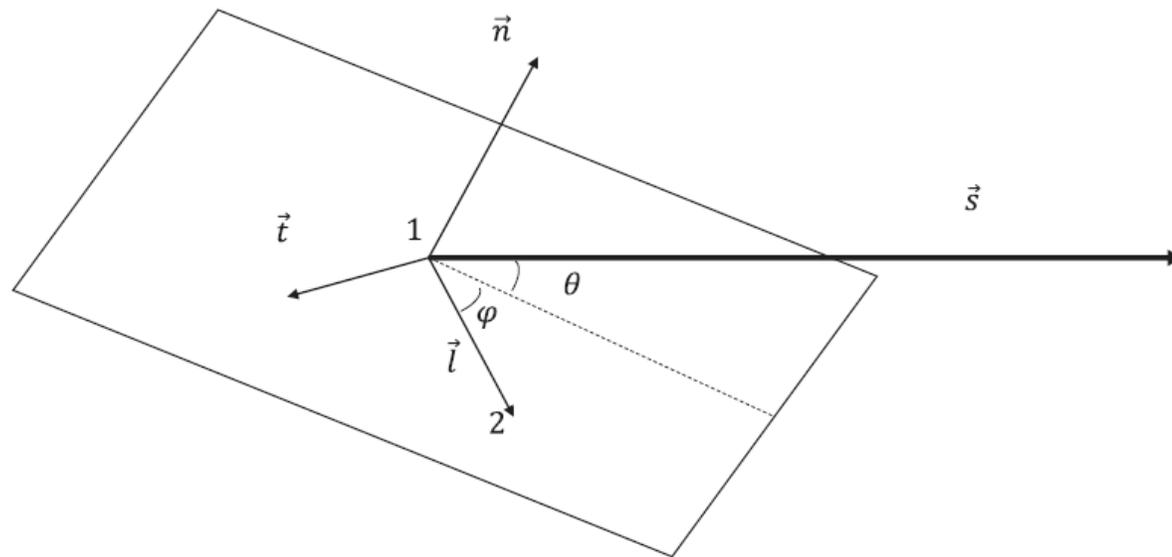
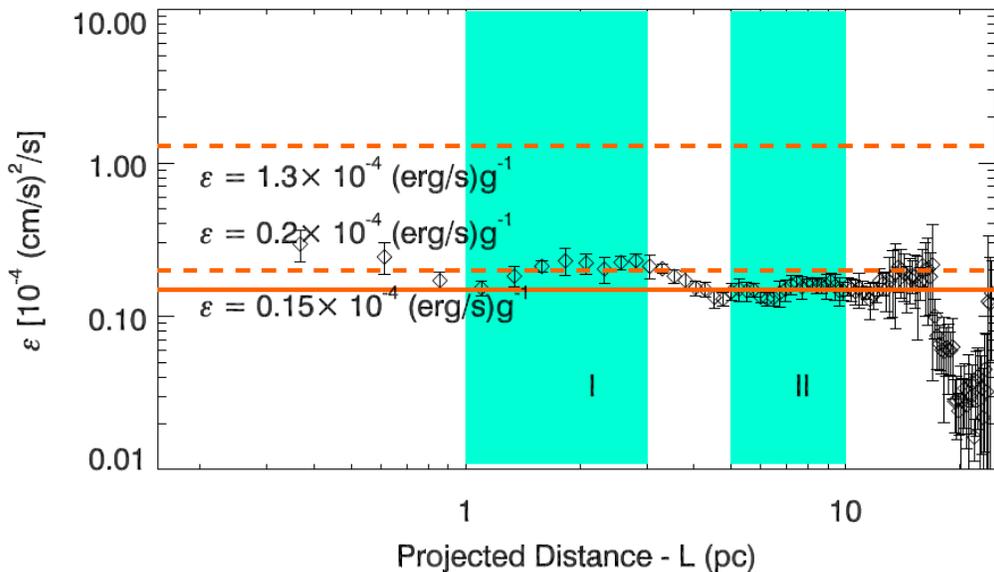
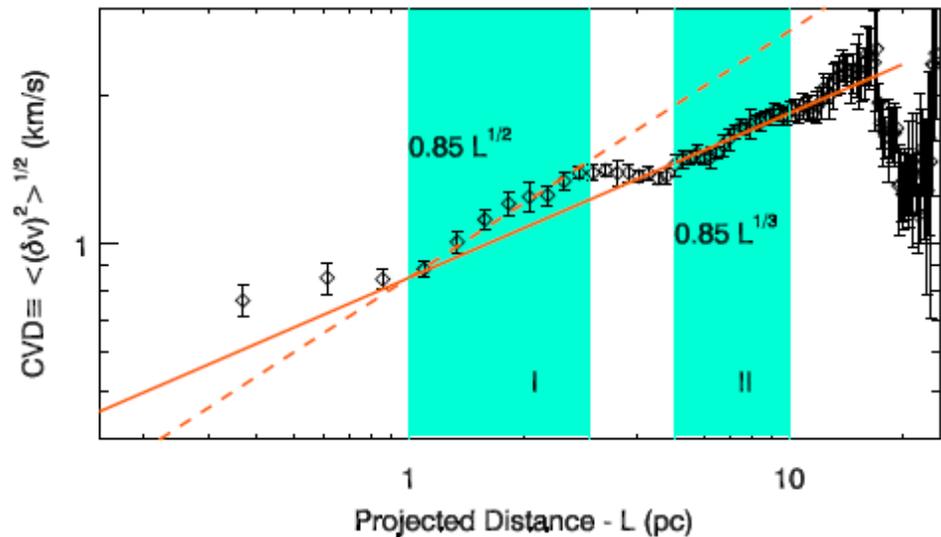
分子云厚度和湍流耗散率



分子云厚度和湍流耗散率



分子云厚度和湍流耗散率



$$CVD^2 = \frac{1}{2} S_{tt}^2 \left(1 - \frac{1}{8} \cos^2 \theta \right) R^{2/3}$$

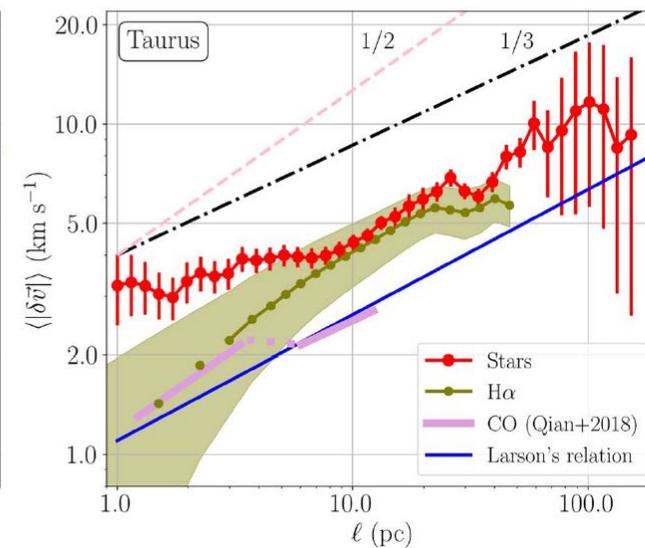
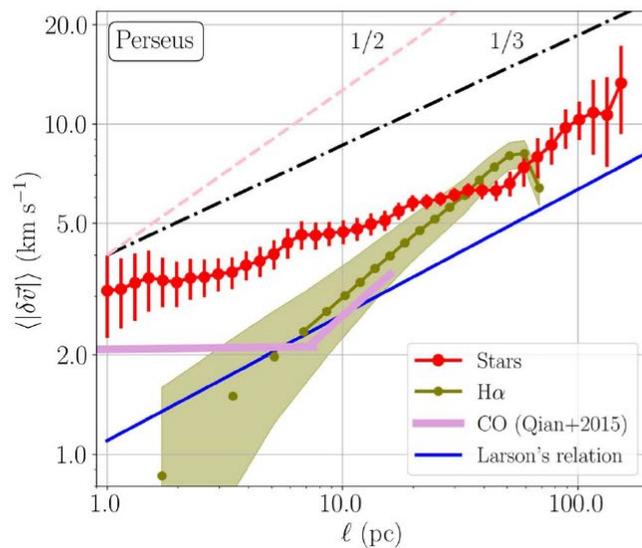
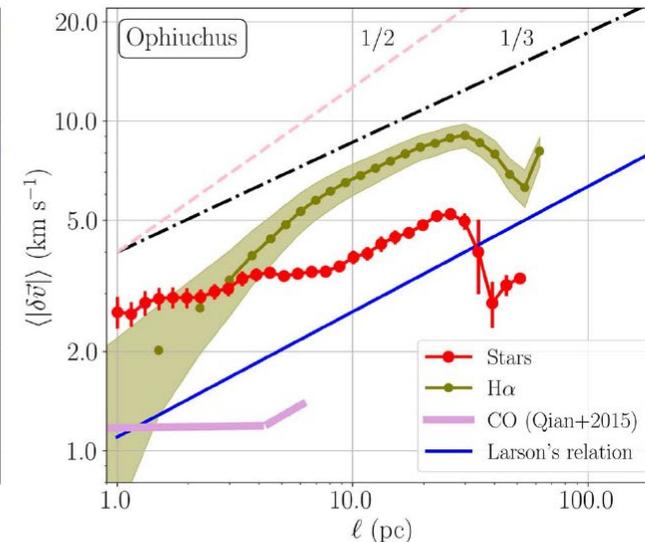
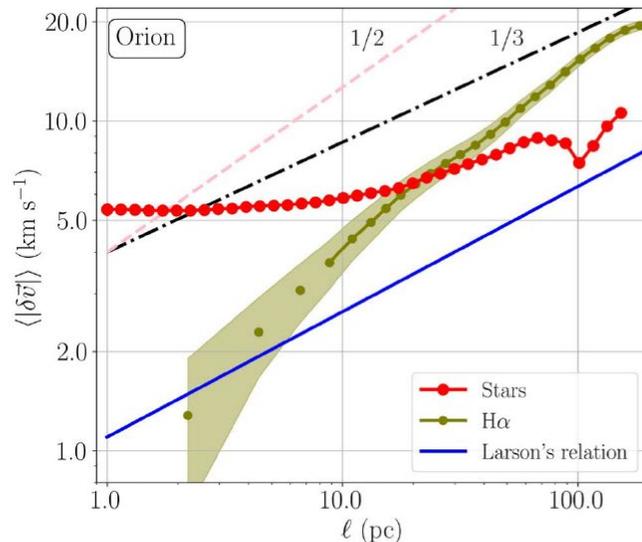
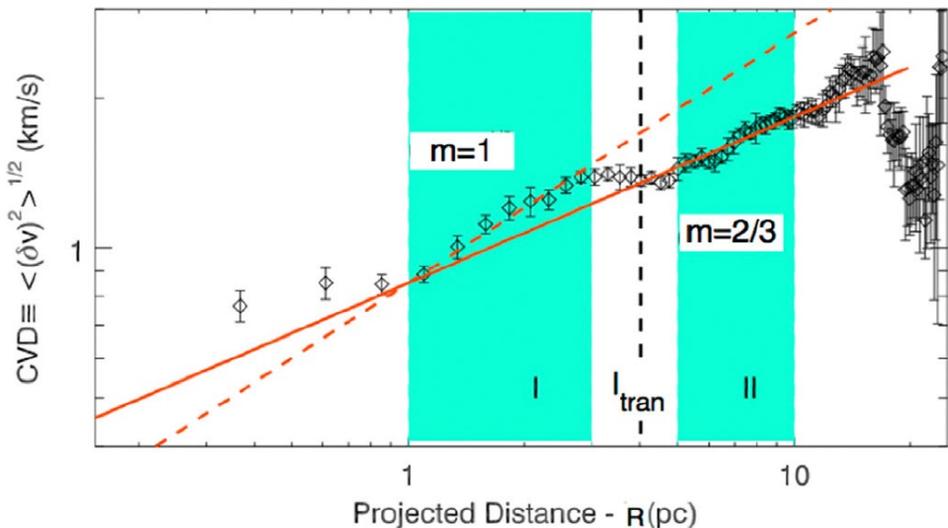
Qian, Li, Gao et al., 2018, ApJ, 864, 116

钱磊, 2021, 天文学报, 62, 7

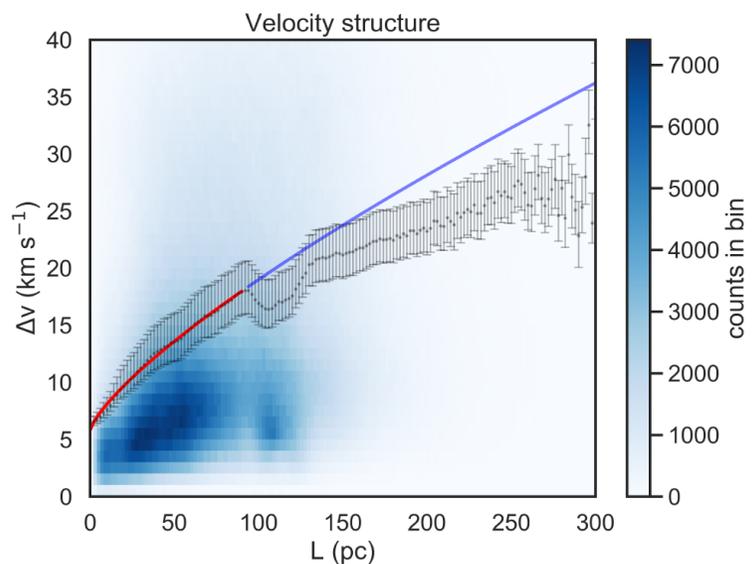
04

总结和展望

总结和展望



Xu, 2020, MNRAS, 492, 1044



Zhang et al. in preparation

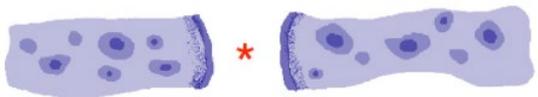
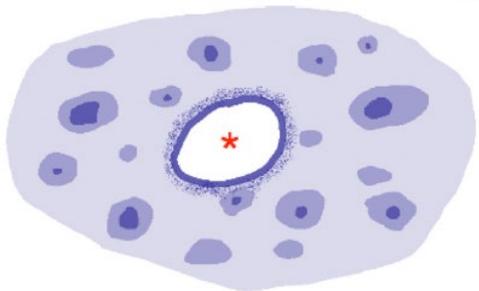
Ha et al., 2022, ApJ, 934, 7

A large radio telescope dish is shown in a dark, starry night sky. The dish is illuminated from below, and several tall, thin towers are visible around it. Blue, energy-like streaks are visible in the sky, suggesting a signal or data being received. The overall scene is dark and atmospheric.

请各位老师批评指正

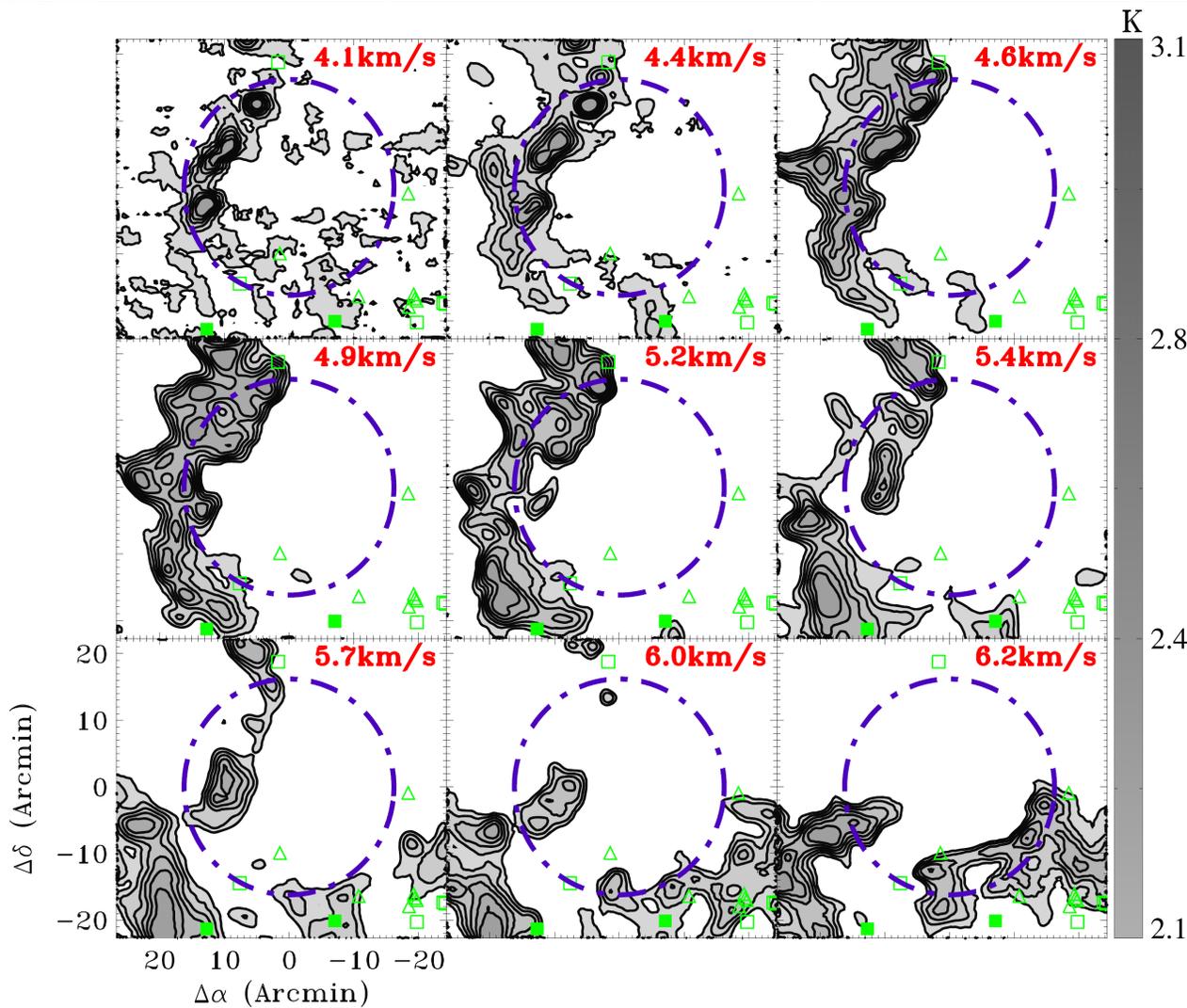
- 谢谢 -

分子云结构和湍流运动

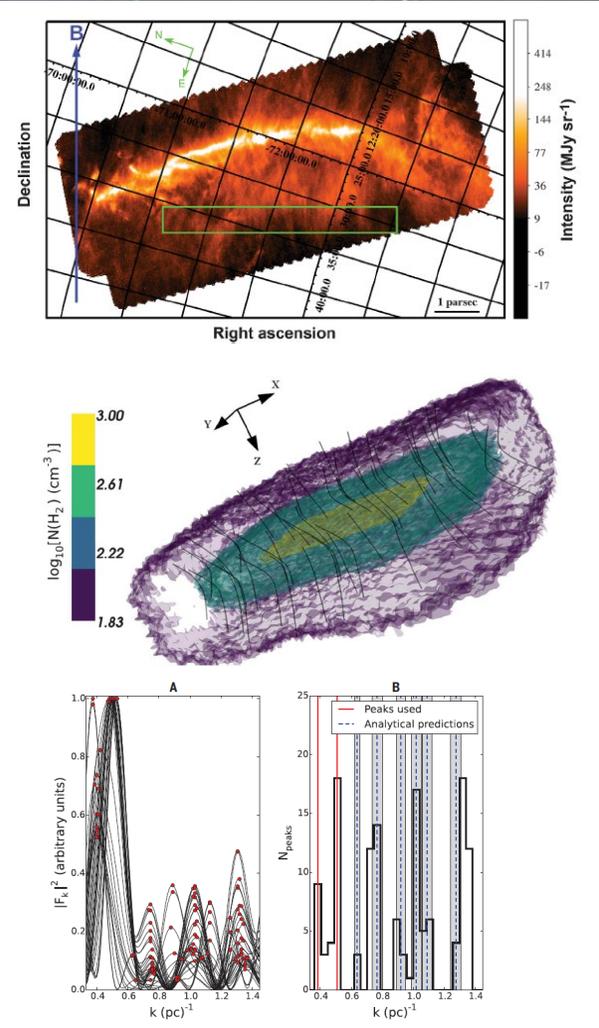


Beaumont, Williams 2010,

ApJ, 709, 791



Li et al. 2015, ApJS, 219, 20



Tritsis, Tassis 2018, Science, 360, 635